

Article

Potentials for the Lanczos Spintensor in Gödel Spacetime

J. López-Bonilla*, R. López-Vázquez & S. Vidal-Beltrán

ESIME-Zacatenco, Instituto Politécnico Nacional, CDMX, México

Abstract

For the Gödel geometry we obtain several potentials for the Lanczos generator associated with the Weyl tensor.

Keywords: Lanczos potential, Gödel’s metric, conformal tensor.

1. Introduction

The Lanczos potential K_{abc} has the properties [1, 2]:

$$K_{abc} = -K_{bac}, \quad K_{abc} + K_{bca} + K_{cab} = 0, \quad K_a^b{}_b = 0, \quad (1)$$

and for the Gödel’s geometry [3-9]:

$$ds^2 = (dx^0)^2 + 2 e^{x^3} dx^0 dx^1 + \frac{1}{2} e^{2x^3} (dx^1)^2 - (dx^2)^2 - (dx^3)^2, \quad (2)$$

we have the expression [10-16]:

$$S_{abc} \equiv K_{abc} + i^* K_{abc} = \frac{i}{9} [U_{ab} m_c + V_{ab} \bar{m}_c - M_{ab} (l_c + n_c)], \quad (3)$$

with the canonical null tetrad [17]:

$$\begin{aligned} (l^a) &= \frac{1}{\sqrt{2}}(1, 0, -1, 0), & (n^a) &= \frac{1}{\sqrt{2}}(1, 0, 1, 0), & (m^a) &= \left(1, -e^{-x^3}, 0, \frac{i}{\sqrt{2}}\right), \\ (\bar{m}^a) &= \left(1, -e^{-x^3}, 0, -\frac{i}{\sqrt{2}}\right), \end{aligned} \quad (4)$$

and:

$$V_{ab} = l_a x m_b, \quad U_{ab} = \bar{m}_a x n_b, \quad M_{ab} = m_a x \bar{m}_b + n_a x l_b. \quad (5)$$

The corresponding expression for the conformal tensor is given by [6, 18]:

*Correspondence: J. López-Bonilla. ESIME-Zacatenco, Instituto Politécnico Nacional, Edif. 5, 1er. Piso, Col. Lindavista CP 07738, CDMX, México. Email: jlopezb@ipn.mx

$$S_{abcr} \equiv C_{abcr} + i {}^*C_{abcr} = -\frac{1}{3}(M_{ab} M_{cr} + V_{ab} U_{cr} + U_{ab} V_{cr}), \quad (6)$$

with the following non-zero Newman-Penrose (NP) quantities [19-21]:

$$\mu = \rho = 2\gamma = 2\varepsilon = \sqrt{2} \alpha = \sqrt{2} \beta = \frac{i}{2}, \quad \psi_2 = -\frac{1}{6}, \quad \phi_{00} = \phi_{22} = 2\phi_{11} = \frac{1}{4}, \quad R = -1, \quad (7)$$

and from (3) it is clear that the NP components of the Lanczos generator are:

$$\Omega_1 = \Omega_6 = -\frac{i}{18}, \quad \Omega_r = 0, \quad r \neq 1, 6, \quad (8)$$

which gives the following components different to zero:

$$\begin{aligned} K_{031} = K_{103} = \frac{1}{2} K_{130} = \frac{1}{18} e^{x^3}, \quad K_{131} = \frac{1}{6} e^{2x^3}, \\ {}^*K_{233} = \frac{1}{2} {}^*K_{200} = \frac{\sqrt{2}}{18}, \quad {}^*K_{201} = {}^*K_{210} = \frac{\sqrt{2}}{9} e^{x^3}, \quad {}^*K_{211} = \frac{5\sqrt{2}}{36} e^{2x^3}, \\ R_{0303} = \frac{1}{2}, \quad R_{0313} = \frac{1}{2} e^{x^3}, \quad R_{1313} = 3 R_{0101} = \frac{3}{4} e^{2x^3}, \quad R_{00} = -1, \quad R_{01} = -e^{x^3}, \quad R_{11} = -e^{2x^3}, \\ C_{0303} = -C_{2323} = -\frac{1}{2} C_{0202} = \frac{1}{6}, \quad C_{0313} = \frac{1}{2} C_{1220} = \frac{1}{6} e^{x^3}, \quad C_{0101} = \frac{1}{4} C_{1313} = -\frac{1}{5} C_{1212} = \frac{e^{2x^3}}{12}, \\ (e_{(0)}^a) \equiv \left(\frac{1}{\sqrt{2}} (l^a + n^a) \right) = (1, 0, 0, 0), \quad (e_{(1)}^a) \equiv \left(\frac{1}{\sqrt{2}} (m^a + \bar{m}^a) \right) = (\sqrt{2}, -\sqrt{2} e^{-x^3}, 0, 0), \\ (e_{(2)}^a) \equiv \left(\frac{i}{\sqrt{2}} (m^a - \bar{m}^a) \right) = (0, 0, 0, -1), \quad (e_{(3)}^a) \equiv \left(\frac{1}{\sqrt{2}} (l^a - n^a) \right) = (0, 0, -1, 0), \quad \Gamma^3_{11} = \frac{e^{2x^3}}{2} \\ \Gamma_{013} = \Gamma_{103} = -\Gamma_{301} = \frac{1}{2} e^{x^3}, \quad \Gamma_{113} = -\Gamma_{311} = \frac{1}{2} e^{2x^3}, \quad \Gamma^3_{01} = \Gamma^0_{13} = \frac{1}{2} e^{x^3}, \quad \Gamma^0_{03} = 1, \\ \Gamma^1_{03} = -e^{-x^3}. \end{aligned}$$

We note that $(e_{(0)}^b)$ is a Killing time-like vector and $(e_{(3)}^b)$ is a constant space-like vector, that is:

$$e_{(0) a;b} + e_{(0) b;a} = 0, \quad e_{(3) a;b} = 0. \quad (10)$$

In Sec. 2 we employ all above information to construct potentials for the Lanczos spintensor in according with the fundamental relation [1]:

$$C_{abcd} = K_{abc;d} - K_{abd;c} + K_{cda;b} - K_{cdb;a} + K_{ad} g_{bc} - K_{ac} g_{bd} + K_{bc} g_{ad} - K_{bd} g_{ac}, \quad (11)$$

where:

$$K^a_b = K_b^a = K^{ac}_{b;c}, \quad K^{abc}_{;c} = 0, \quad (12)$$

with the non-zero components:

$$-\frac{1}{2} K^0_0 = K^1_1 = K^3_3 = \frac{1}{6}, \quad K^0_1 = -\frac{1}{2} e^{x^3}. \quad (13)$$

2. Several generators for the Lanczos potential

Now we study if the following relation has solution for some constant q :

$$q K^{abc} = C^{abc r}_{;r}, \quad (14)$$

which is equivalent to a set of eight differential equations in the NP formalism:

$$\begin{aligned} q \Omega_0 &= D\psi_1 - \bar{\delta}\psi_0 + (4\alpha - \pi) \psi_0 - 2(2\rho + \varepsilon) \psi_1 + 3\kappa \psi_2, \\ q \Omega_1 &= D\psi_2 - \bar{\delta}\psi_1 + \lambda \psi_0 + 2(\alpha - \pi) \psi_1 - 3\rho \psi_2 + 2\kappa \psi_3, \\ q \Omega_2 &= D\psi_3 - \bar{\delta}\psi_2 + 2\lambda \psi_1 - 3\pi \psi_2 + 2(\varepsilon - \rho) \psi_3 + \kappa \psi_4, \\ q \Omega_3 &= D\psi_4 - \bar{\delta}\psi_3 + 3\lambda \psi_2 - 2(2\pi + \alpha) \psi_3 + (4\varepsilon - \rho) \psi_4, \\ q \Omega_4 &= \delta\psi_1 - \Delta\psi_0 + (4\gamma - \mu) \psi_0 - 2(2\tau + \beta) \psi_1 + 3\sigma \psi_2, \\ q \Omega_5 &= \delta\psi_2 - \Delta\psi_1 + \nu \psi_0 + 2(\gamma - \mu) \psi_1 - 3\tau \psi_2 + 2\sigma \psi_3, \\ q \Omega_6 &= \delta\psi_3 - \Delta\psi_2 + 2\nu \psi_1 - 3\mu \psi_2 + 2(\beta - \tau) \psi_3 + \sigma \psi_4, \\ q \Omega_7 &= \delta\psi_4 - \Delta\psi_3 + 3\nu \psi_2 - 2(2\mu + \gamma) \psi_3 + (4\beta - \tau) \psi_4, \end{aligned} \quad (15)$$

where we can employ the values (7) and (8) to obtain $q = -\frac{9}{2}$, thus (14) implies the expression:

$$K^{abc} = -\frac{2}{9} C^{abcr}{}_{;r}, \quad (16)$$

which may be considered as the inversion of (11), that is, the Weyl tensor is a generator of the Lanczos potential in the Gödel's cosmological model.

The next stage is to search solutions for the tensor equation:

$$K_{abc} = P_{ca;b} - P_{cb;a}, \quad P_{jr} = P_{rj}, \quad P = P^r{}_r, \quad (17)$$

via its corresponding NP equations:

$$\begin{aligned} \Omega_0 &= 2[D\Omega_{01} - \delta\Omega_{00} + (2\bar{\alpha} + 2\beta - \bar{\pi})\Omega_{00} - 2(\bar{\rho} + \varepsilon)\Omega_{01} - 2\sigma\Omega_{10} + \bar{\kappa}\Omega_{02} + 2\kappa\Omega_{11}], \\ \Omega_1 &= 2\left[D\Omega_{11} - \delta\Omega_{10} + \frac{1}{8}DP + \mu\Omega_{00} + (2\bar{\alpha} - \bar{\pi})\Omega_{10} - \pi\Omega_{01} - \sigma\Omega_{20} - 2\bar{\rho}\Omega_{11} + \kappa\Omega_{21} + \bar{\kappa}\Omega_{12}\right], \\ \Omega_2 &= 2\left[\bar{\delta}\Omega_{11} - \Delta\Omega_{10} - \frac{1}{8}\bar{\delta}P + \nu\Omega_{00} + (2\bar{\gamma} - \bar{\mu})\Omega_{10} - \lambda\Omega_{01} - \tau\Omega_{20} - 2\bar{\tau}\Omega_{11} + \rho\Omega_{21} + \sigma\Omega_{12}\right], \\ \Omega_3 &= 2[\bar{\delta}\Omega_{21} - \Delta\Omega_{20} + 2\nu\Omega_{10} + (2\bar{\gamma} - 2\gamma - \bar{\mu})\Omega_{20} - 2\lambda\Omega_{11} + 2(\alpha - \bar{\tau})\Omega_{21} + \bar{\sigma}\Omega_{22}], \\ \Omega_4 &= 2[D\Omega_{02} - \delta\Omega_{01} + \bar{\lambda}\Omega_{00} + 2(\beta - \bar{\pi})\Omega_{01} + (2\bar{\varepsilon} - 2\varepsilon - \bar{\rho})\Omega_{02} - 2\sigma\Omega_{11} + 2\kappa\Omega_{12}], \\ \Omega_5 &= 2\left[D\Omega_{12} - \delta\Omega_{11} + \frac{1}{8}\delta P + \mu\Omega_{01} + \bar{\lambda}\Omega_{10} - \pi\Omega_{02} - 2\bar{\pi}\Omega_{11} + (2\bar{\varepsilon} - \bar{\rho})\Omega_{12} - \sigma\Omega_{21} + \kappa\Omega_{22}\right], \\ \Omega_6 &= 2\left[\bar{\delta}\Omega_{12} - \Delta\Omega_{11} - \frac{1}{8}\bar{\delta}P + \nu\Omega_{01} + \bar{\nu}\Omega_{10} - \lambda\Omega_{02} - 2\bar{\mu}\Omega_{11} + (2\bar{\beta} - \bar{\tau})\Omega_{12} - \tau\Omega_{21} + \rho\Omega_{22}\right], \\ \Omega_7 &= 2[\bar{\delta}\Omega_{22} - \Delta\Omega_{21} + \bar{\nu}\Omega_{20} + 2\nu\Omega_{11} - 2(\bar{\mu} + \gamma)\Omega_{21} - 2\lambda\Omega_{12} + (2\alpha + 2\bar{\beta} - \bar{\tau})\Omega_{22}], \\ \Omega_1 + \bar{\Omega}_1 &= 2\left[D\Omega_{11} - \Delta\Omega_{00} - \frac{1}{8}DP + 2(\gamma + \bar{\gamma})\Omega_{00} - (2\bar{\tau} + \pi)\Omega_{01} - (2\tau + \bar{\pi})\Omega_{10} + \bar{\kappa}\Omega_{12} + \kappa\Omega_{21}\right], \end{aligned} \quad (18)$$

$$\bar{\Omega}_2 - \Omega_5 = 2 \left[\delta \Omega_{11} - \bar{\delta} \Omega_{02} + \frac{1}{8} \delta P + (2\bar{\mu} - \mu) \Omega_{01} - \bar{\lambda} \Omega_{10} + 2(\alpha - \bar{\beta}) \Omega_{02} + (\bar{\rho} - 2\rho) \Omega_{12} + \sigma \Omega_{21} \right],$$

$$\Omega_6 + \bar{\Omega}_6 = 2 \left[D \Omega_{22} - \Delta \Omega_{11} + \frac{1}{8} DP + \nu \Omega_{01} + \bar{\nu} \Omega_{10} - (\bar{\tau} + 2\pi) \Omega_{12} - (\tau + 2\bar{\pi}) \Omega_{21} + 2(\varepsilon + \bar{\varepsilon}) \Omega_{22} \right],$$

where $\Omega_{ab} = \bar{\Omega}_{ba}$ are the NP components of $Q_{ab} = P_{ab} - \frac{P}{4} g_{ab}$. It is easy to see that a solution of (18) is $P_{ab} = -\frac{1}{9} R_{ab}$, that is, $\Omega_{ab} = -\frac{1}{9} \phi_{ab}$ and $P = -\frac{1}{9} R = \frac{1}{9}$, therefore (17) gives the relation:

$$K_{abc} = -\frac{1}{9} (R_{ca;b} - R_{cb;a}), \quad (19)$$

which is compatible with (16) due to the Bianchi identities [5, 6, 21]; hence the Ricci tensor also is a generator of the Lanczos spintensor in Gödel spacetime.

With (9) is possible to deduce the wave equation for some quantities of interest, for example:

$$\square K^{abc} \equiv K^{abc;r}{}_{;r} = -3 K^{abc}, \quad \square R_{ab} = 2 \square K_{ab} = -6 K_{ab}, \quad \square e_{(0)a} = -e_{(0)a}, \quad (20)$$

besides:

$$K_{ab;c} + K_{bc;a} + K_{ca;b} = 0, \quad K^{ab}{}_{;b} = 0, \quad R_{abcr} K^{abc} = 0, \\ R^r{}_{[a} K_{b]cr} - R^r{}_c K_{abr} - R^r{}_q K_{r[a}{}^q g_{b]c} = K_{abc}, \quad (21)$$

$$K_{ab} = -\frac{1}{6} (2 e_{(0)a} e_{(0)b} + e_{(1)a} e_{(1)b} + e_{(2)a} e_{(2)b}) = \frac{1}{6} (R_{ab} - 2R_{acrb} G^{cr}),$$

and K_{ab} turns out to be a potential for the Lanczos tensor:

$$K_{abc} = -\frac{2}{9} (K_{ca;b} - K_{cb;a}), \quad (22)$$

thus (22) is the inversion of (12).

An adequate linear combination of R_{ab} and K_{ac} gives an interesting generator:

$$K_{abc} = -\frac{\sqrt{2}}{18} (B_{ca;b} - B_{cb;a}), \quad B^{ac}{}_{;c} = 0, \quad (23)$$

$$B_{ac} = \frac{1}{\sqrt{2}} (6K_{ac} - R_{ac}) = \frac{1}{\sqrt{2}} (2R_{ac} + g_{ac} + e_{(3)a} e_{(3)c}) = \sqrt{2} (G_{ac} + \frac{1}{2} e_{(3)a} e_{(3)c}),$$

$$\square B_{ac} = -6\sqrt{2} K_{ac},$$

where $G_{ab} = R_{ab} - \frac{R}{2}g_{ab}$ is the Einstein tensor, and it is remarkable the fulfillment of the Gauss equation:

$$R_{abcr} = B_{ac} B_{br} - B_{ar} B_{bc}, \tag{24}$$

$$B_{ac} = -\sqrt{2} R_{abrc} G^{br}, \quad *R^{*abcr} R_{abcr} = 0,$$

hence B_{ac} generates the Lanczos potential via the differential relation (23) and also the curvature tensor through the algebraic expression (24).

Finally, in [22] were obtained the following properties in Gödel geometry:

$$K_{abc} = \frac{\sqrt{2}}{3} *C_{abcr} e_{(3)}^r, \quad *K_{abc} = -\frac{\sqrt{2}}{3} C_{abcr} e_{(3)}^r, \tag{25}$$

which indicate an algebraic connection between Lanczos and Weyl tensors, unlike the differential relationship (16).

Received December 12, 2018; Accepted January 25, 2019

References

1. C. Lanczos, *The splitting of the Riemann tensor*, Rev. Mod. Phys. **34**, No. 3 (1962) 379-389.
2. J. D. Zund, *Sur le spineur de Lanczos en relativité générale*, Comptes Rendus Acad. Sci. (Paris) **A276** (1973) 1629-1631.
3. K. Gödel, *An example of a new type of cosmological solution of Einstein's field equations of gravitation*, Rev. Mod. Phys. **21**, No. 3 (1949) 447-450.
4. S. Chandrasekhar, J. P. Wright, *The geodesics in Gödel's Universe*, Proc. Nat. Acad. Sci. USA **47** (1961) 341-347.
5. J. L. Synge, *Relativity: the general theory*, North-Holland, Amsterdam (1976).
6. H. Stephani, D. Kramer, M. MacCallum, C. Hoenselaers, E. Herlt, *Exact solutions of Einstein's field equations*, Cambridge University Press (2003).
7. J. López-Bonilla, G. Ovando, J. Rivera, *Embedding of the Gödel metric*, Aligarh Bull. Math. **17** (1997-98) 63-66.
8. Z. Ahsan, J. López-Bonilla, A. Rangel-Merino, *Wave equation in Gödel model*, J. Vect. Rel. **4**, No. 2 (2009) 80-81.
9. C. Hernández-Aguilar, A. Domínguez-Pacheco, J. López-Bonilla, B. Thapa, *Embedding of the Gödel cosmological model*, The SciTech, J. of Sci. & Tech. **4**, No. 1 (2015) 18-20.
10. J. López-Bonilla, G. Ovando, *Lanczos spintensor for the Gödel metric*, Gen. Rel. Grav. **31**, No. 7 (1999) 1071-1074.
11. V. Gaftoi, J. López-Bonilla, G. Ovando, *Lanczos potential for the Gödel cosmological model*, Czech. J. Phys. **52**, No. 6 (2002) 811-813.
12. R. García-Olivo, J. López-Bonilla, S. Vidal-Beltrán, *Gödel's geometry: Embedding and Lanczos spintensor*, EJTP **3**, No. 12 (2006) 55-58.
13. J. H. Caltenco, R. Linares, J. López-Bonilla, *A potential for conformal and curvature tensors in Gödel spacetime*, South East Asian J. Math. & Math. Sc. **5**, No. 1 (2006) 77-79.

14. Z. Ahsan, V. Barrera-Figueroa, J. López-Bonilla, *A potential for Gödel cosmological model*, Bull. Calcutta Math. Soc. **99**, No. 4 (2007) 323-332.
15. Z. Ahsan, J. H. Caltenco, J. López-Bonilla, *Lanczos potential for the Gödel spacetime*, Ann. der Phys. **16**, No. 4 (2007) 311-313.
16. Z. Ahsan, J. H. Caltenco, R. Linares, J. López-Bonilla, *Lanczos generator in Gödel geometry*, Comm. in Phys. **20**, No. 1 (2010) 9-14.
17. J. López-Bonilla, R. López-Vázquez, J. Morales, G. Ovando, *Petrov types and their canonical tetrads*, Prespacetime Journal **7**, No. 8 (2016) 1176-1186.
18. J. López-Bonilla, R. López-Vázquez, J. Morales, G. Ovando, *Maxwell, Lanczos, and Weyl spinors*, Prespacetime Journal **6**, No. 6 (2015) 509-520.
19. J. López-Bonilla, R. López-Vázquez, J. Morales, G. Ovando, *Spin coefficients formalism*, Prespacetime Journal **6**, No. 8 (2015) 697-709.
20. A. H. Hasmani, P. I. Andharia, *Algebraic computations of spin coefficients in Newman-Penrose formalism using Mathematica*, J. of Dynamical Systems & Geometric Theories **9**, No. 1 (2011) 2736.
21. P. Lam-Estrada, J. López-Bonilla, R. López-Vázquez, A. K. Rathie, *Newman-Penrose equations, Bianchi identities, and Weyl-Lanczos relations*, Prespacetime Journal **6**, No. 8 (2015) 684-696.
22. I. Guerrero-Moreno, J. López-Bonilla, R. López-Vázquez, S. Vidal-Beltrán, *Lanczos generator in terms of the conformal tensor in Gödel and type D vacuum geometries*, American-Eurasian J. of Sci. Res. **13**, No. 4 (2018) 67-70